

# GalSim

Galaxy Simulation Software Package

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LSST Simulations Meeting

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<https://github.com/GalSim-developers/GalSim>

# Overview

- Simulates images of galaxies, stars, noise, etc.
- Originally designed for Great3 WL challenge.
- Can draw using FFTs, photon shooting, or real space convolution/rendering as appropriate.
- Was designed to have shapes accurate to  $2e-4$ .
- Has a variety of options for all aspects of the rendering (PSF, galaxy profile, WCS, noise, etc.)

# Code Design

- Open source.
- Modular structure (easy to add new features).
- Ability to trade between speed and accuracy.
- Intuitive Python layer wrapping fast C++.
- Even more intuitive configuration file interface.
- All code reviewed by typically 2 team members.

# PSF Profiles

- Gaussian
- Moffat
- Kolmogorov
- Airy (with optional central obscuration)
- OpticalPSF (incl. Zernike to arbitrary order, obscuration, struts, or arbitrary spider pattern)
- InterpolatedImage (arbitrary image, e.g. PSFEx)

# Galaxy Profiles

- Exponential
- DeVaucouleurs
- Sersic
- InterpolatedImage (can use any provided image as a starting profile)
- RealGalaxy (several available HST catalogs, properly deconvolving HST PSF)

# Transformations

- Any profile can be:
  - sheared
  - dilated
  - rotated
  - shifted
  - scaled in flux
  - added to another profile
  - convolved with another profile
  - deconvolved by another profile (FFT only!)

# Weak Lensing

- Realistic shear models
  - [NFWHalo](#) model
  - Cosmological (or other) [PowerSpectrum](#)
  - Can specify E and B power spectra separately
- These also include the corresponding convergence field so magnification and shear can be done consistently.

# Random Numbers

- Can draw random numbers from:
  - `UniformDeviate` samples from uniform distribution
  - `GaussianDeviate` samples from normal distribution
  - `PoissonDeviate` handles Poisson number counts
  - other more exotic distributions (binomial,  $\chi^2$ , Weibull, gamma)
  - `DistDeviate` can use any arbitrary (user-provided) probability distribution
- "Random" numbers are deterministic given a seed, even when splitting job into multiple sub-jobs.



# World Coordinate Systems

- Handles a variety of WCS transformations
  - `PixelScale` (the simple case of square pixels)
  - `JacobianWCS` (arbitrary local Jacobian defined by  $du/dx$ ,  $du/dy$ ,  $dv/dx$ ,  $dv/dy$ )
  - `AffineTransform` (Jacobian with an arbitrary offset)
  - `FitsWCS` (read standard WCS types from FITS file)
  - `UVFunction` (any arbitrary function from  $(x,y)$  to tangent plane  $(u,v)$ )
  - `RaDecFunction` (any arbitrary function from  $(x,y)$  to  $(RA, Dec)$ )

# Rendering Images

- Most calculations are deferred until image is rendered.
- Can use either FFT or photon shooting
  - Caveat: Photon shooting doesn't work when deconvolving. This includes RealGalaxy profiles.
- Can draw at arbitrary position in image.
- Can place postage stamps in tiles or randomly.
- Can set parameters that trade off accuracy vs speed.

# Chromatic Profiles

- **Chromatic** is simplest case of a profile with a uniform SED.
- **ChromaticAtmosphere** will apply the correct shift and shear for DCR to a given (e.g. Kolmogorov) profile
- **ChromaticAiry** properly handles the  $\lambda/D$  parameter across the bandpass.
- **ChromaticOpticalPSF** properly scales the Zernike coefficients with wavelength (in addition to  $\lambda/D$ ).
- Renders images by integrating over **Bandpass**.

# Noise Models

- GaussianNoise
- PoissonNoise
- CCDNoise includes Poisson photon noise (with optional gain) and Gaussian read noise.
- VariableGaussianNoise (sigma different for each pixel)
- CorrelatedNoise
- Can also whiten images to remove existing correlated noise.

# Detector Effects

- Brighter-fatter effect
  - Currently has Gruen implementation of Antilogus et al
  - Craig Lage is working on incorporating his physical model tracking individual electrons.
- Non-linearity
- Edge distortion and tree rings could be modeled with custom WCS, but nothing native currently.

# Shape Measurement

- Includes `HSM` module for measuring shapes of galaxies.
- `FindAdaptiveMom` performs an adaptive moments measurement
- `EstimateShear` accounts for PSF convolution
  - KSB method
  - Re-Gaussianization (Hirata & Seljak, 2003)
  - "Linear" (H&S correction to Bernstein & Jarvis, 2002 method)

# Galsim Demos

- Demo scripts are in Galsim/examples directory.
  - demo1.py ... demo13.py
  - demo1.yaml ... demo11.yaml
- After installing Galsim, you can run (for example)
  - python demo1.py
  - galsim demo1.yaml
- Most features are in at least one demo script.
- Python and YAML versions produce identical output files.

## Demo1.py

```
import galsim

gal = galsim.Gaussian(flux=1.e5, sigma=2)

psf = galsim.Gaussian(flux=1., sigma=1.)

final = galsim.Convolve(gal, psf)

image = final.drawImage(scale=0.2)

noise = galsim.GaussianNoise(sigma=30.)
image.addNoise(noise)

file_name = 'output/demo1.fits'
image.write(file_name)
```

## Demo1.yaml

```
gal :
  type : Gaussian
  sigma : 2
  flux : 1.e5

psf :
  type : Gaussian
  sigma : 1

image :
  pixel_scale : 0.2
  noise :
    type : Gaussian
    sigma : 30

output :
  dir : output_yaml
  file_name : demo1.fits
```



SAOImage ds9

File Edit View Frame Bin Zoom Scale Color Region WCS Analysis Help

File:

Object:

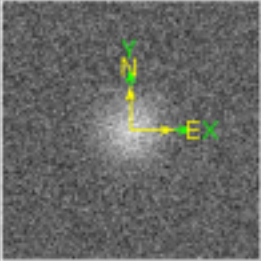

Value:

WCS:

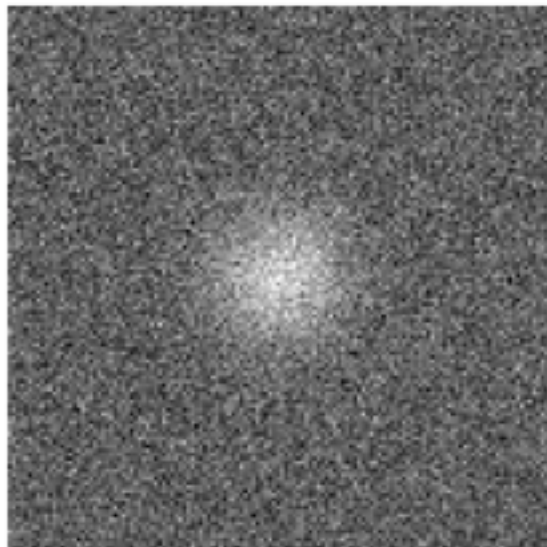
Physical X:  Y:

Image X:  Y:

Frame 1 x:   °

file	edit	view	frame	bin	zoom	scale	color	region	wcs	help
-	+	to fit	zoom 1/8	zoom 1/4	zoom 1/2	zoom 1	zoom 2	zoom 4	zoom 8	



# Demo11.yaml

```
eval_variables :
  fpixel_scale : &pixel_scale 0.2
  atheta : &theta 0.17 degrees
  ftel_diam : &tel_diam 4
  fexp_time : &exp_time 300
  fimage_size : &image_size 1800
  inobjects : &nobjects 288

psf :
  type : InterpolatedImage
  image : "data/example_sdss_psf_sky0.fits.bz2"
  scale : *pixel_scale

gal :
  type : COSMOSGalaxy
  gal_type :
    type : List
    items : [ 'parametric', 'real' ]
    index : { type : RandomBinomial, N : 1, p : 0.3 }
  noise_pad_size : 11.3
  index : { type : Random }
  shear : { type : PowerSpectrumShear }
  magnification : { type : PowerSpectrumMagnification }
  rotation : { type : Random }
  scale_flux :
    type : Eval
    str : "(tel_diam**2 / (2.4**2*(1.-0.33**2))) * exp_time"

stamp :
  draw_method : no_pixel
```

```
image :
  type : Scattered
  size : *image_size
  nobjects : *nobjects
  index_convention : 0
  noise :
    type : Gaussian
    variance : 5.0e4 # Total variance including whatever the symmetrize process needs.
    symmetrize : 8
wcs :
  type : Tan
  dudx : { type : Eval, str : 'math.cos(theta.rad()) * pixel_scale' }
  dudy : { type : Eval, str : '-math.sin(theta.rad()) * pixel_scale' }
  dvdx : { type : Eval, str : 'math.sin(theta.rad()) * pixel_scale' }
  dvdy : { type : Eval, str : 'math.cos(theta.rad()) * pixel_scale' }
  units : arcsec
  origin : center
  ra : 19.3 hours
  dec : -33.1 degrees
random_seed : 24783923
nproc : -1

input :
  cosmos_catalog :
    dir : "data"
    file_name : real_galaxy_catalog_example.fits
    use_real : True
  power_spectrum :
    e_power_function : "data/cosmo-fid.zmed1.00.out"
    units : radians
    grid_spacing : 90 # arcsec

output :
  dir : output_yaml
  file_name : tabulated_power_spectrum.fits.fz
```

File Edit View Frame Bin Zoom Scale Color Region WCS Analysis Help

File tabulated\_power\_spectrum.fits.fz

Object

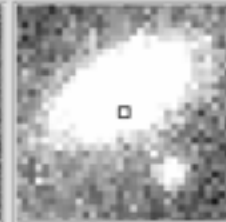
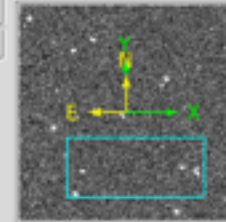
Value 3747.35

fk5  $\alpha$  19:17:50.537  $\delta$  -33:07:32.73

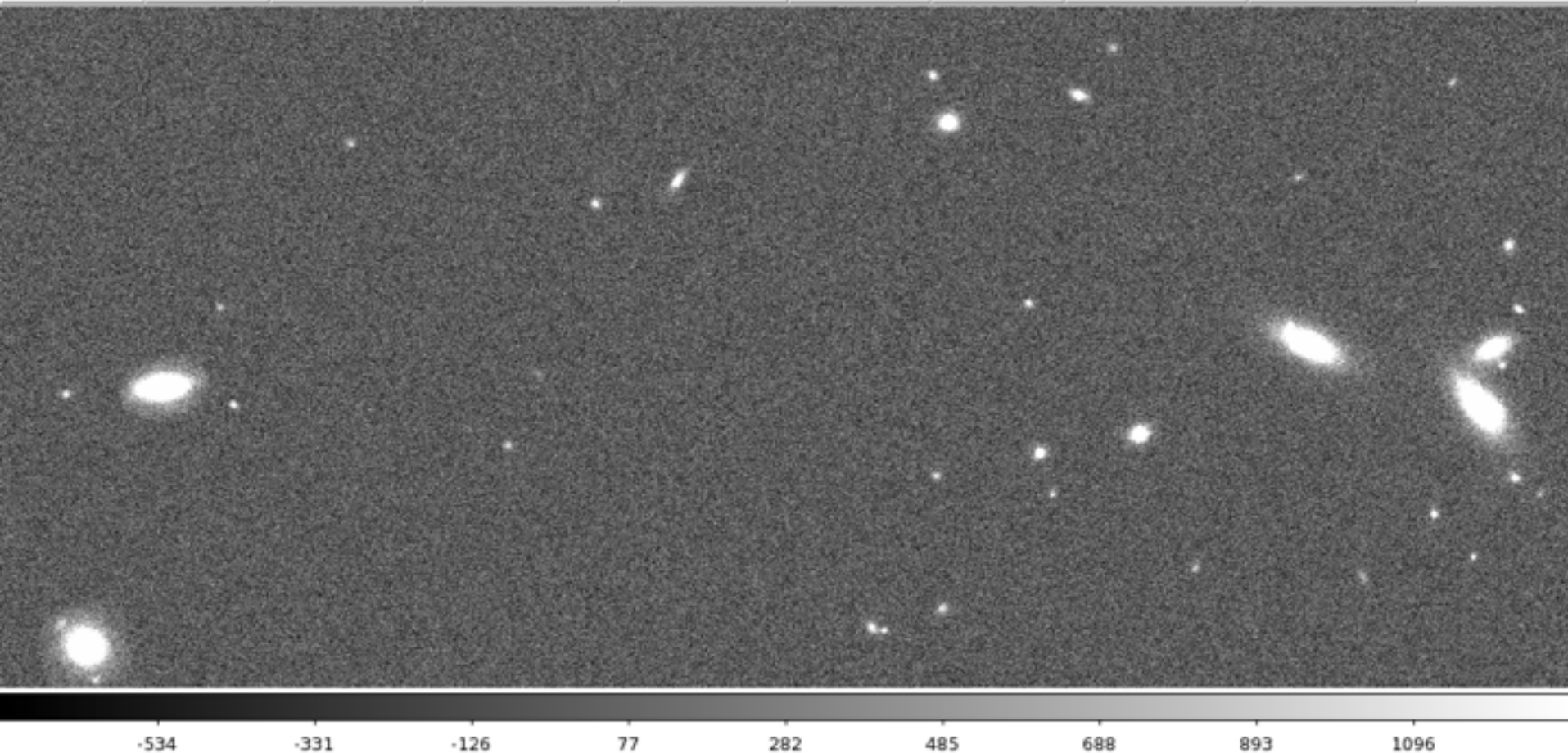
Physical X 1493.500 Y 435.000

Image X 1493.500 Y 435.000

Frame 1 x 1.000 0.000 \*



file	edit	view	frame	bin	zoom	scale	color	region	wcs	help
linear	log	power	square root	squared	asinh	sinh	histogram	min max	zscale	



-534

-331

-126

77

282

485

688

893

1096

# LSST-Specific Features

- LSST WCS
  - Includes knowledge of chip layout as well as expected SIP component of each chip's WCS.
  - Currently in code review, will be in the next release, version 1.4.
  - Requires a LSST-DM stack installation to use.
- Nothing else yet.

# Planned Features

- Chromatic RealGalaxy [#640]
  - Can take images from two or more observed bands and simulate a different band.
- Realistic Atmospheric PSF [#549]
  - Track phase screen from multiple atmospheric layers using correct Fourier optics
- Both are currently in code review.

# Planned Features

- Physics-based models of brighter/fatter [#722]
  - Craig Lage is working on incorporating his physics-based simulation of the CCD silicon to improve the implementation of brighter/fatter.
  - Intrinsically photon-shooting method, but for FFT methods, we essentially draw the image and then photon shoot that.
  - For now, achromatic, but will add chromaticity.

# Planned Features

- Model of LSST PSF [#556]
  - Spider pattern
  - Reference values for aberrations.
  - Approximate rms values of aberrations.
  - Model of how they will vary (correlated?) across the field of view.
  - Election diffusion in CCD



# Planned Features

- Chromatic photon shooting [#540]
- Add chromatic options to config module [#510]
- Proper field dependence of aberrations [#716]
- Vignetting, non-uniform QE [#553]
- Extinction, opacity [#541, #550]
- Realistic galaxy profiles from hydro simulations [#669]

# Planned Features

- Automatically chose FFT for bright objects and photon shooting for faint objects [#209]
- More image artifacts
  - cosmic rays
  - saturation
  - bad/hot pix, columns, etc.
  - cross-talk
  - charge transfer inefficiency
- Lots more: <https://github.com/GalSim-developers/GalSim/issues>