



# ISR Roadmap for the Brighter Fatter Effect in LSSTCam

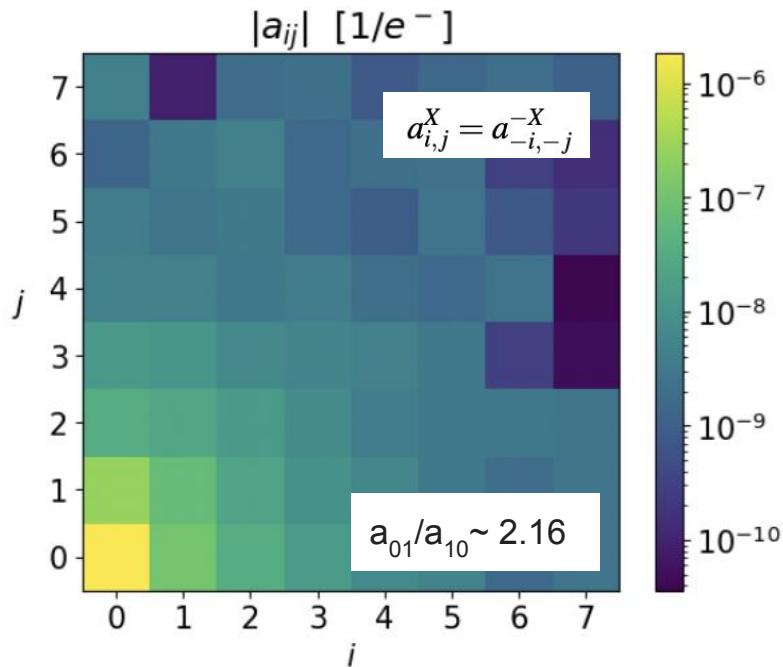
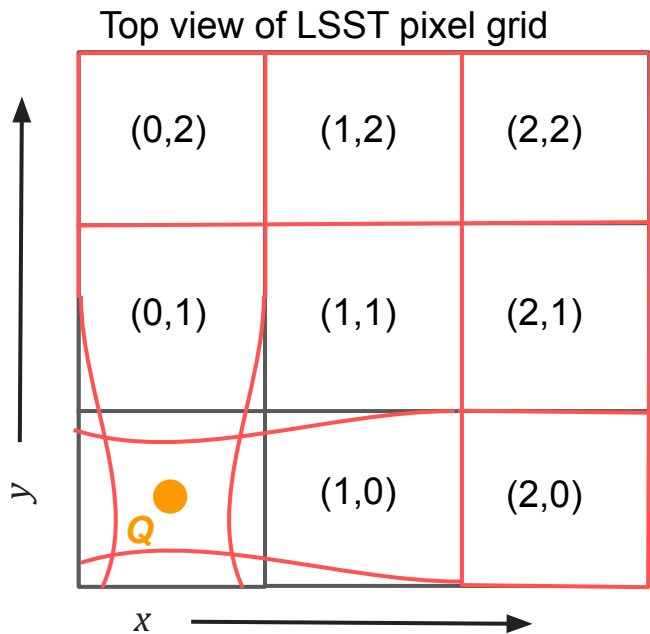


Alex Broughton

Science Pipelines Meeting  
1/31/2024



# Pixels are distorted in LSST Sensors due to charge accumulation

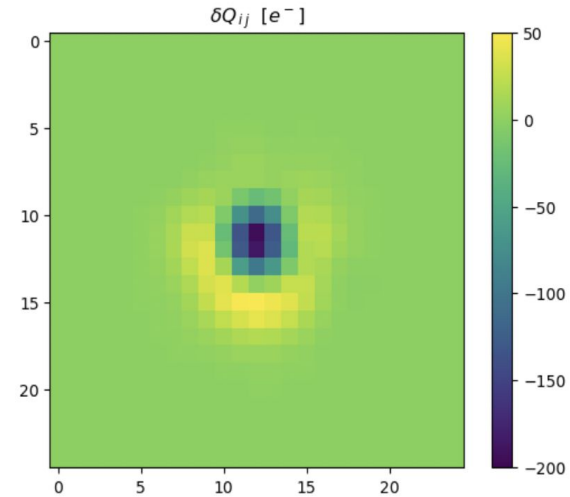
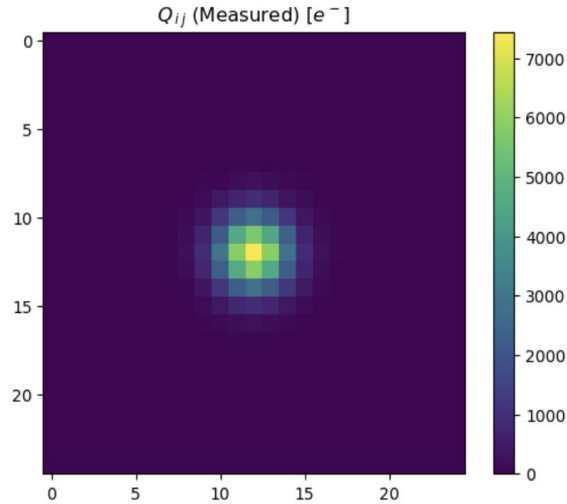


**More than 50% of charge displacement happens beyond 4 pixels away!**

*Broughton et al. (2023)*

# The Brighter-Fatter Effect (BFE)

The BFE makes bright sources appear larger.  
e.g. Calibration stars, brightest cluster galaxies (BCGs), type Ia SNe etc.



These higher-order effects make up 30% of the total effect near sensor saturation

At low signal levels, the pixel-to-pixel effects are approximated well by a constant fractional area change “matrix”

At high signal levels, the pixel-to-pixel effects are non-trivial and need to be measured empirically.

# The Correction

**Calculate Covariances**



**Derive a 2D kernel from covariances**



**Apply to Image**

*Based on*

*Coulton et al. 2018*

*Astier et al. 2019*

*Broughton et al. 2023*

# The Correction

Calculate Covariances



Derive a 2D kernel from covariances



Apply to Image

Pixel-pixel covariances derived from PTC

$$C_{ij}(\mu) = \frac{\mu}{g} \left[ \delta_{i0}\delta_{j0} + a_{ij}\mu g + \frac{2}{3}[\mathbf{a} \otimes \mathbf{a} + \mathbf{ab}]_{ij}(\mu g)^2 + \frac{1}{6}[2\mathbf{a} \otimes \mathbf{a} \otimes \mathbf{a} + 5\mathbf{a} \otimes \mathbf{ab}]_{ij}(\mu g)^3 + \dots \right] + n_{ij}/g^2$$

Original Poisson term      Modified by change in area      Higher-order BFEs

Based on  
*Coulton et al. 2018*  
*Astier et al. 2019*  
*Broughton et al. 2023*

# The Correction

Calculate Covariances



Derive a 2D kernel from covariances



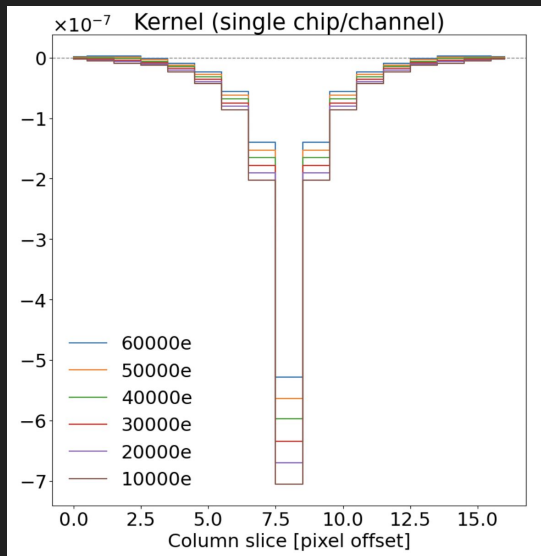
Apply to Image

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$$C(\mathbf{x} - \mathbf{x}') = -\mu^2 \frac{\partial}{\partial x} \cdot \frac{\partial}{\partial x} K$$



Based on

Coulton et al. 2018

Astier et al. 2019

Broughton et al. 2023

# The Correction

Calculate Covariances



Derive a 2D kernel from covariances



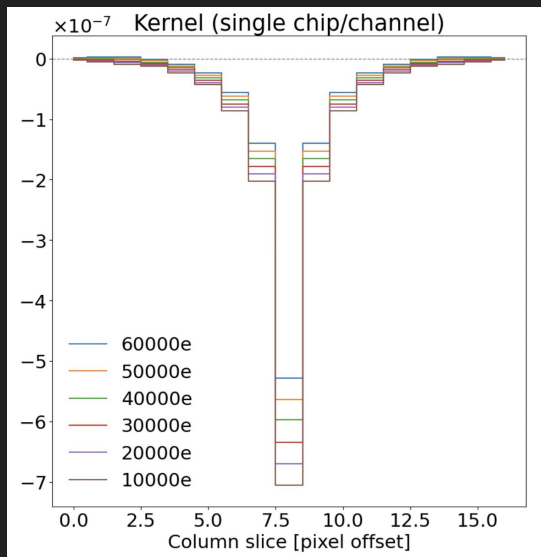
Apply to Image

Pixel-pixel covariances derived from PTC

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Step 1

$$\Phi = F * K$$

Step 2

$$V = F \nabla \Phi$$

Step 3

$$\delta F = \frac{1}{2} \nabla \cdot V$$

$$\hat{F} = F + \delta F$$

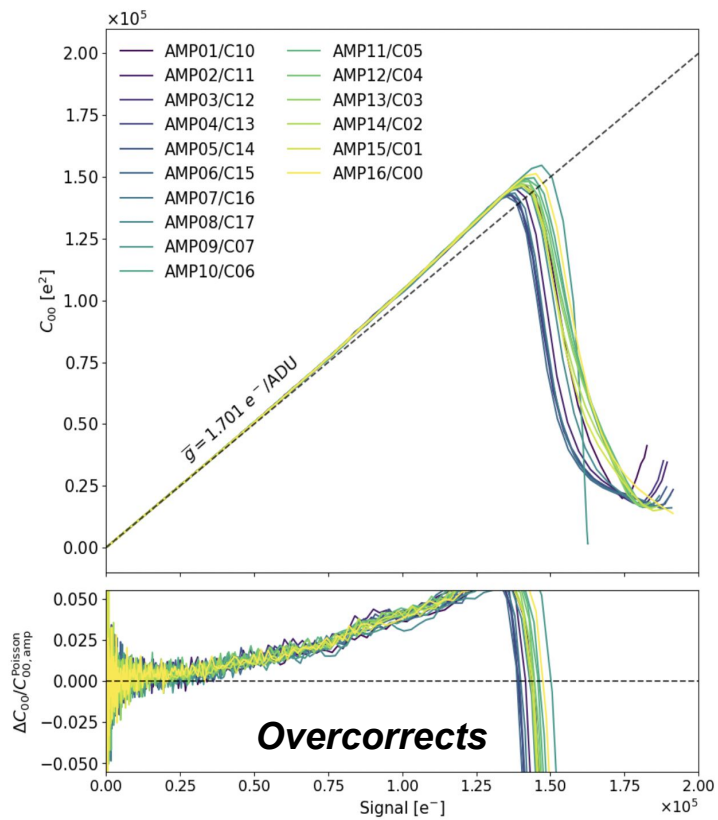
Based on  
 Coulton et al. 2018  
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 Broughton et al. 2023



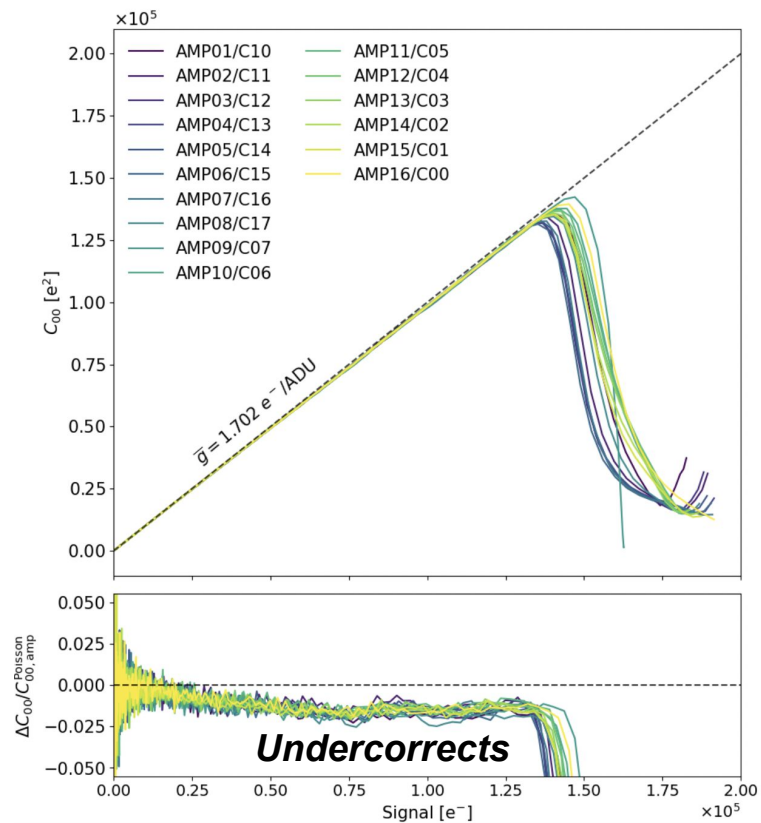
*The correction is derived from measured pixel correlations  
at some arbitrary signal level.*

*Can this correction reconstruct the expected  
variance in flat fields?*

## 10k electrons



## 60k electrons



# 1st improvement to ISR:

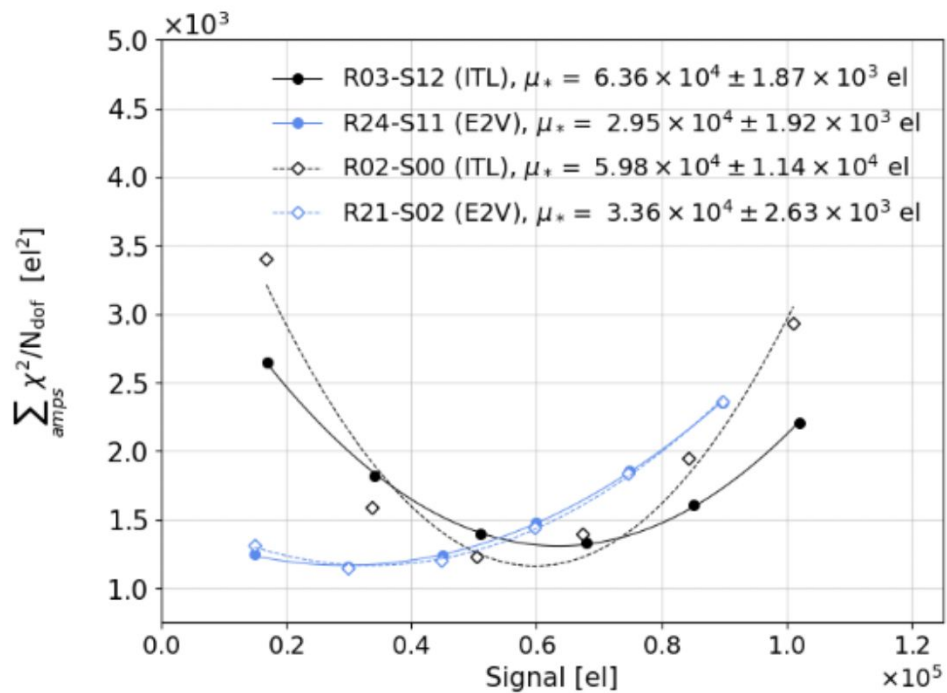
*"Flux-sampling"*

DM-41952

*Let's pick the "sweet spot" signal level that best reconstructs the Poisson form of the PTC.*

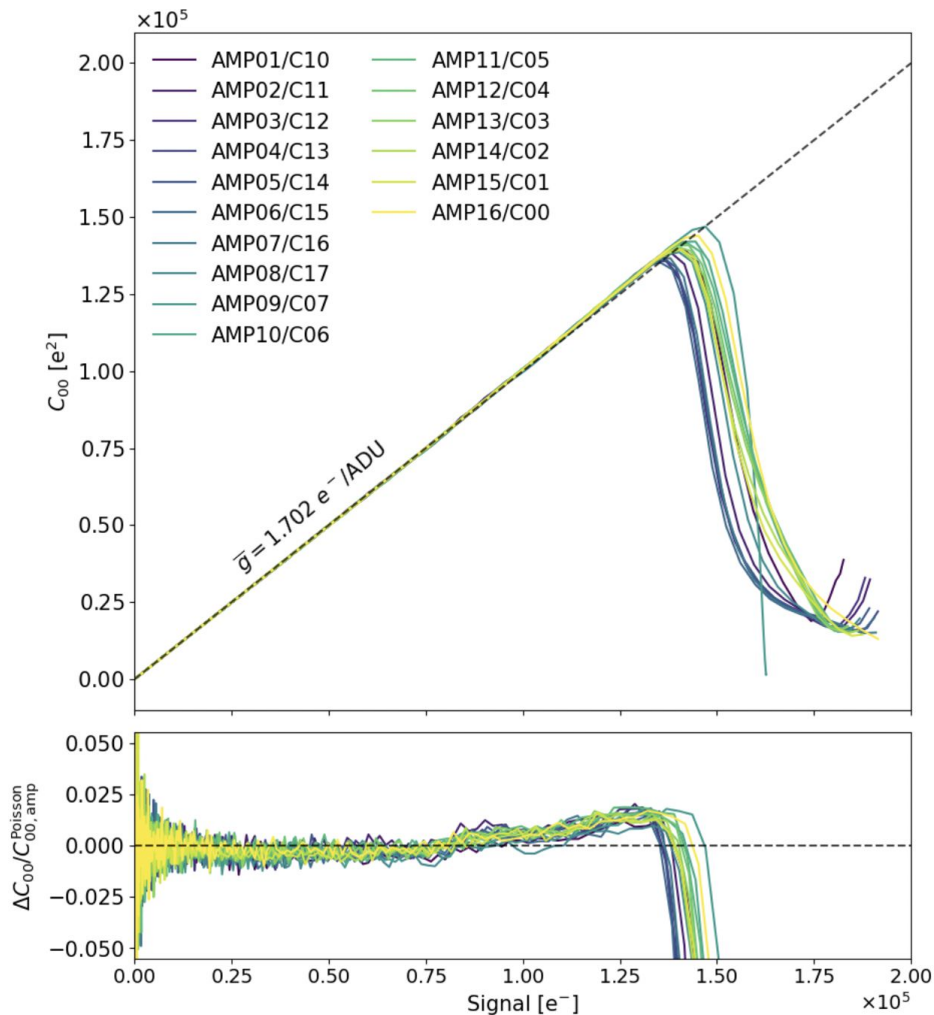
*Test kernels at multiple signal levels, and find the one that minimizes the  $\chi^2$*

$$\chi^2 = \sum_{\mu} (C_{00} - C_{00}^{Poisson})^2 w_{\mu}$$



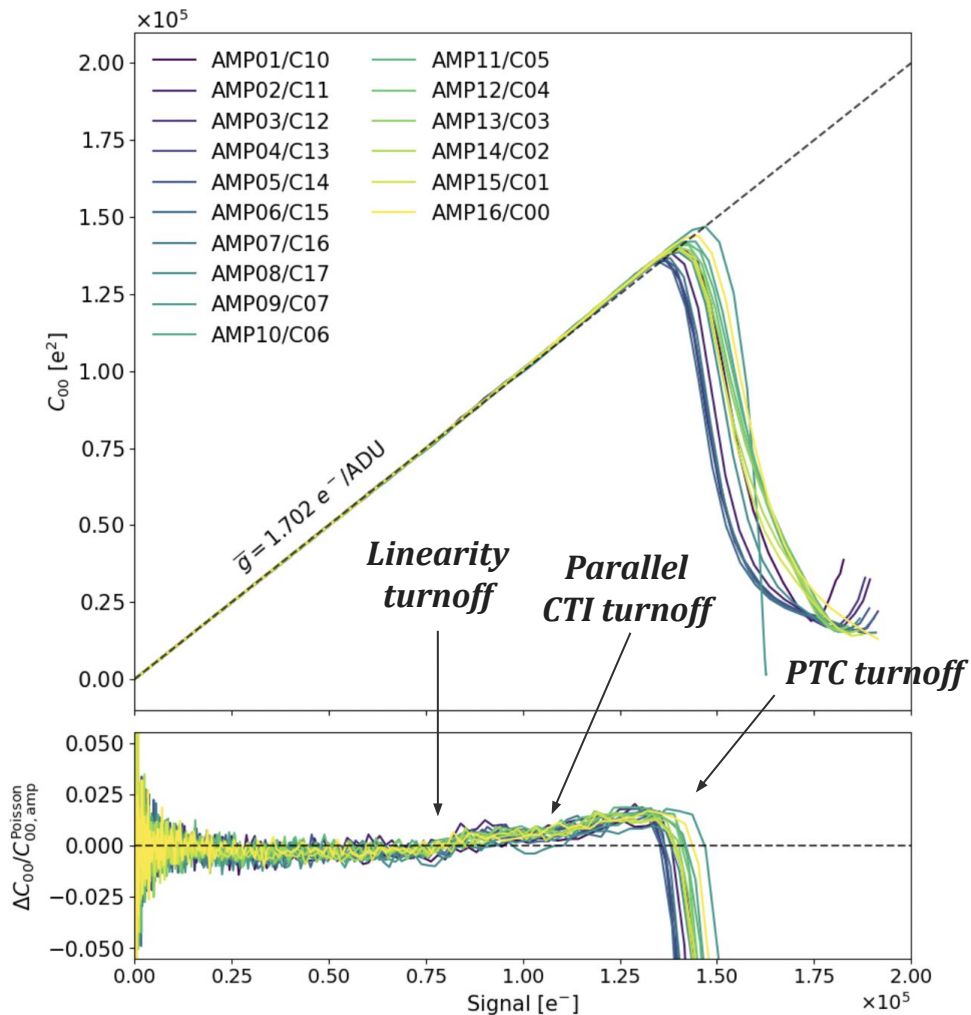
Corrects 94%  
of the effect  
in Flat Fields

Corrects 90%  
of the anisotropy  
between x/y



Corrects 94%  
of the effect  
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Corrects 90%  
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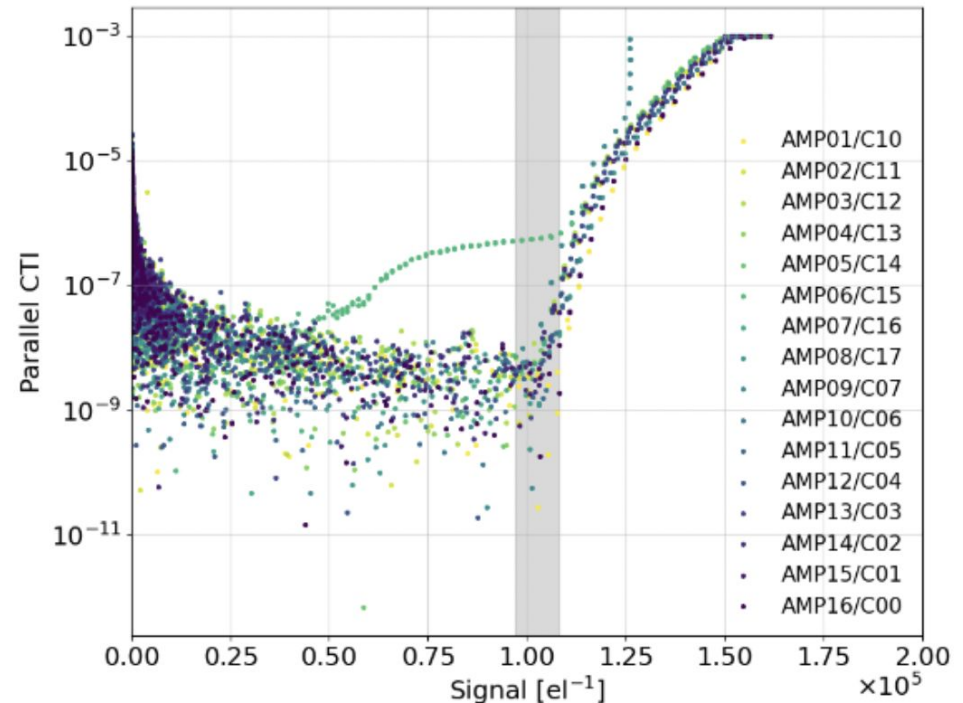


# 2nd improvement to ISR:

*"Parallel CTI turnoff"*

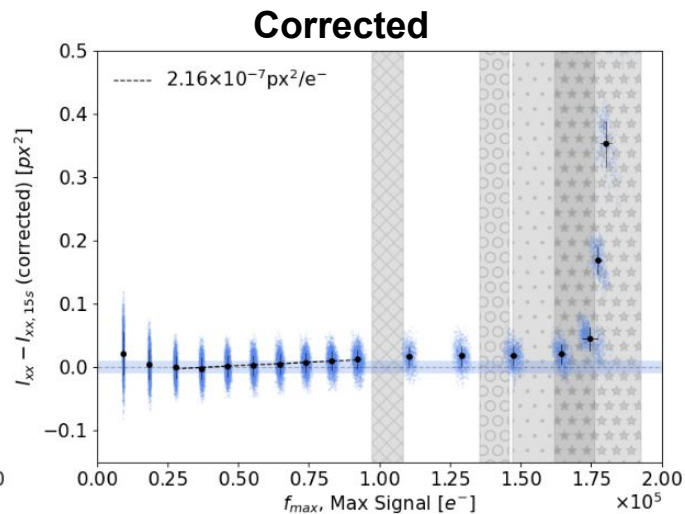
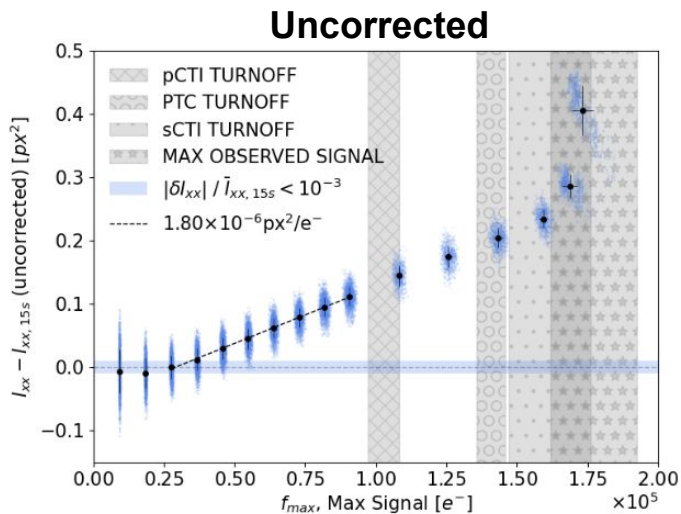
DM-38989

*Defined by point with  
n# of consecutive  
increases in parallel  
CTI above thresholds  
in x and y*



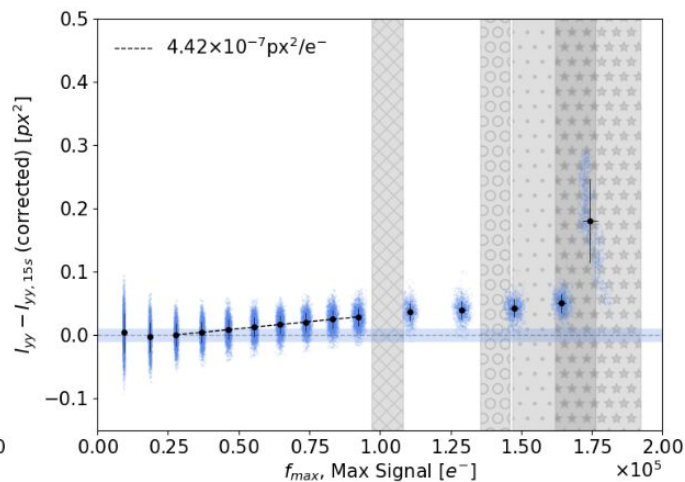
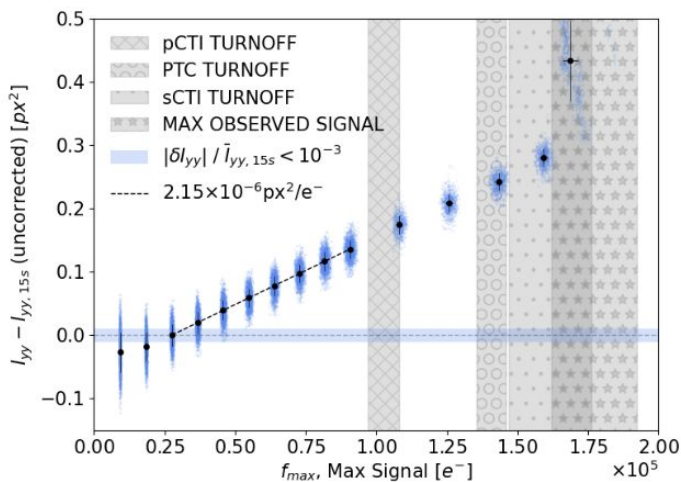
Corrects 90%  
of the effect  
in stars

$I_{xx}$



Corrects  
77% of the  
anisotropy  
between x/y

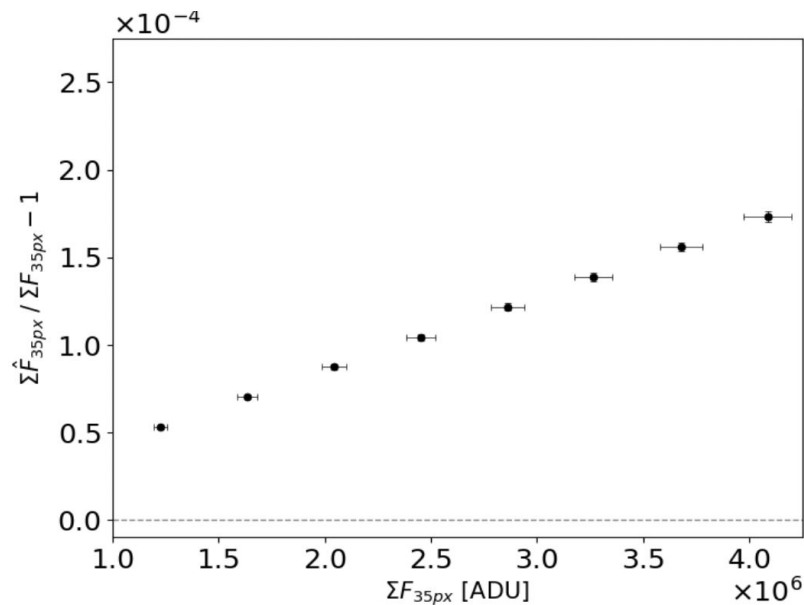
$I_{yy}$





# Why is the overall correction better in flat fields than in stars?

1. *Most of the correction is dominated by  $K_{00}$ , but realistically most of the BFE is contributed by correlations  $> 4\text{px}$  away.*
2. *The application of the correction deviates from Gauss's Law on small scales, resulting in loss of charge conservation in stars (!)*



*More flux is gained by the central pixels than is taken from the neighboring pixels*

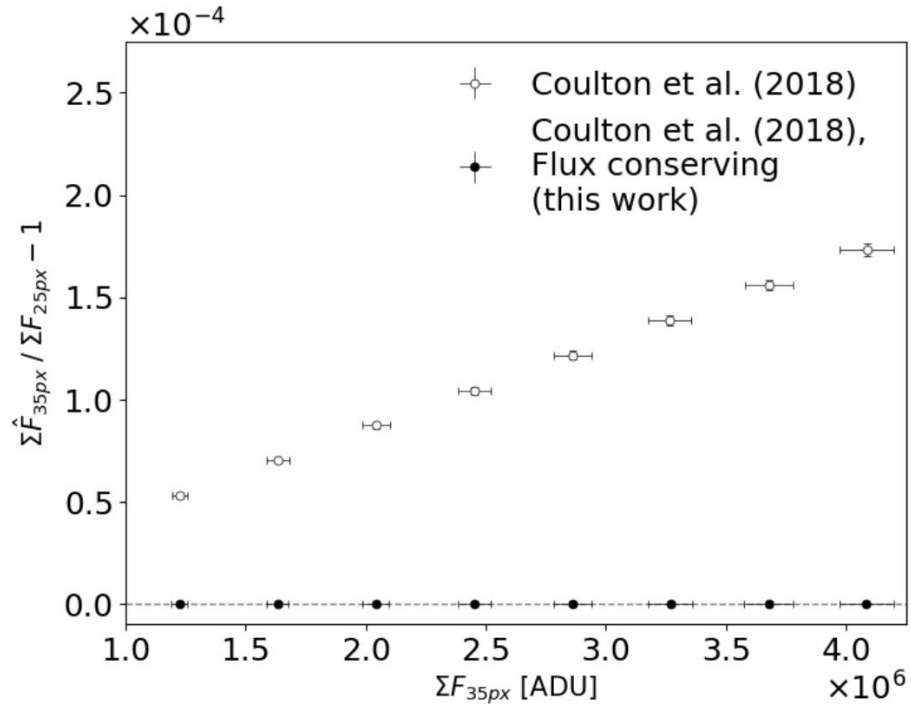
*Flux is conserved, but only in the continuous limit*

$$\langle \delta F \rangle = 0$$

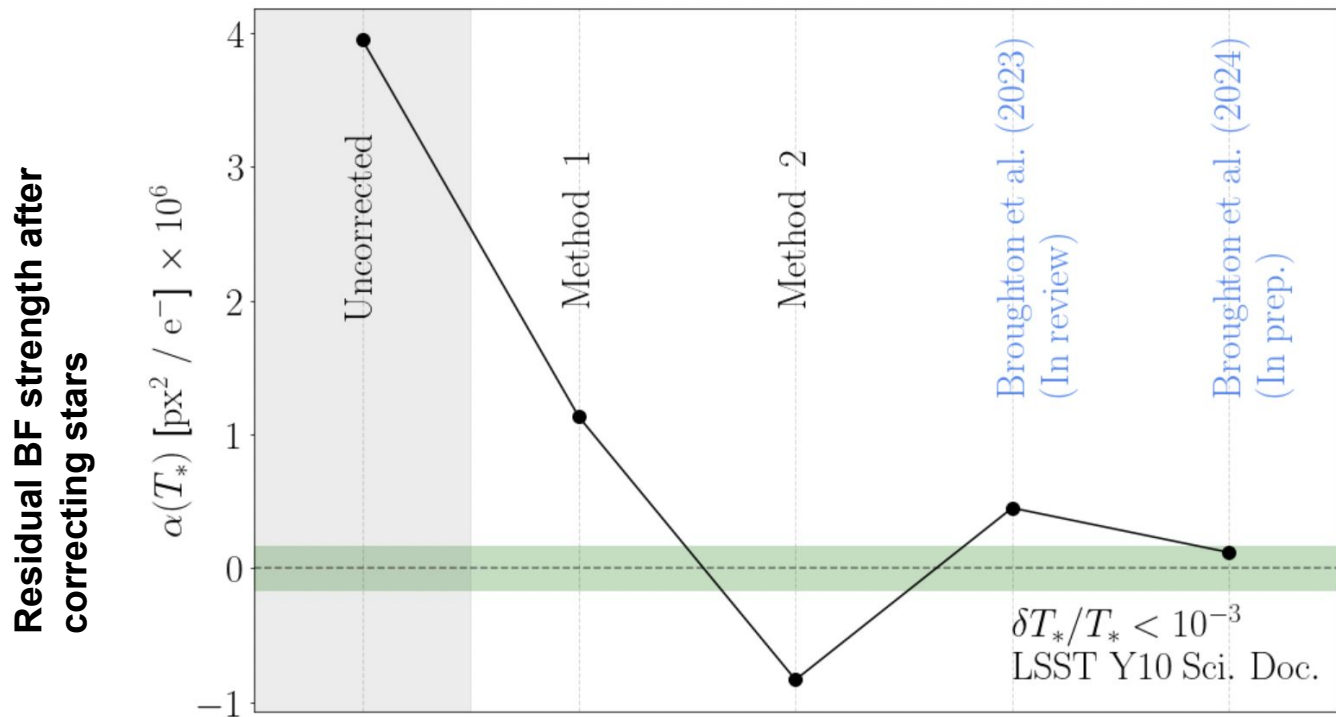
*Poor modeling of local charge transport = worse overall correction*

# 3rd improvement to ISR: *"Flux conserving corrections"*

DM-38555



# Improvements can reconstruct true star size $T_* = \langle I_{xx} + I_{yy} \rangle$



*Method 1:* Using kernel derived from high signal

*Method 2:* Using kernel derived from low signal

*Method 3:* Using kernel at the level that best reconstructs Poisson noise in flat fields.

*Method 4:* Method 3 + flux-conserving corrections (in prep.)

# Why is the anisotropy correction better in flat fields than in stars?

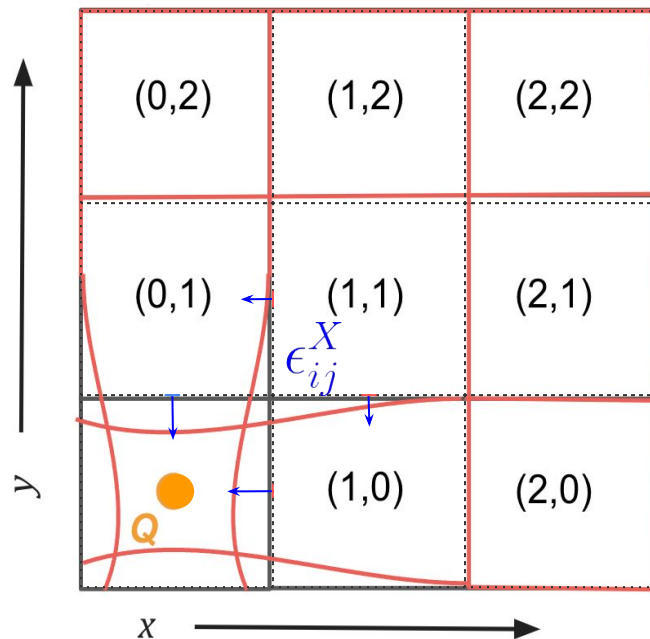
*Poor local modeling of sub-pixel charge transport also due to the assumption that the curl of the displacement field created by the accumulated charges is zero.*

The kernel is only defined by the divergence:

$$\epsilon \propto \nabla_x K$$

... which assumes:

$$\nabla_x \times \epsilon = 0$$



# 4th improvement to ISR:

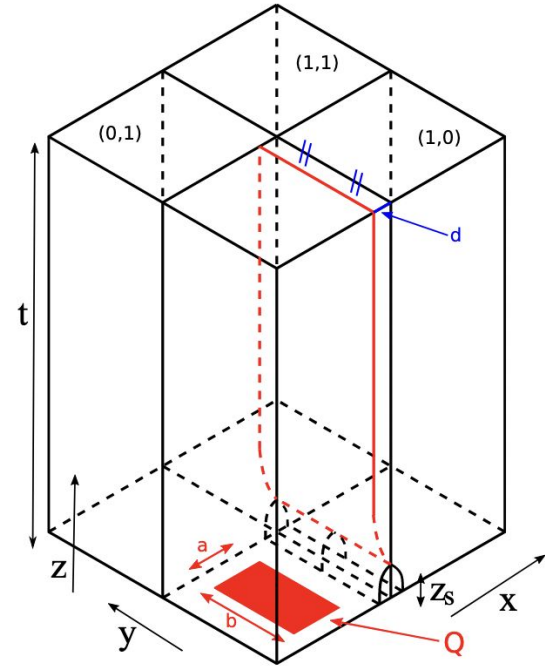
"Adds Astier+23 correction"

DM-39515

Takes in the scalar a-matrix (1 number/pixel) and fits the electrostatic solution for the boundary shifts given a charge +Q in a potential well to derive the vector a-matrix (4 numbers/pixel)

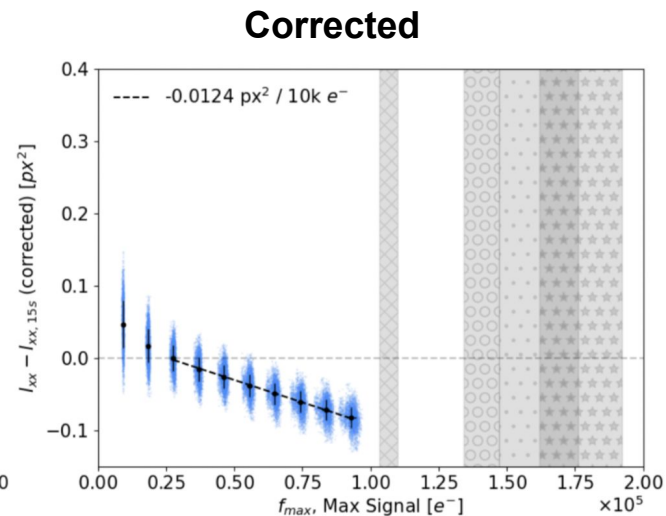
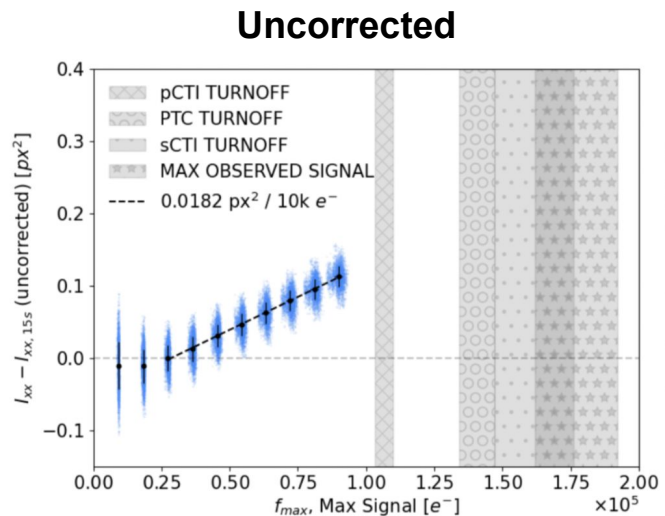
$$\begin{array}{l} \text{e-model} \\ \text{fit} \\ a_{ij} \rightarrow \vec{a}_{ij} = \end{array} \begin{pmatrix} a_{ij}^N \\ a_{ij}^E \\ a_{ij}^S \\ a_{ij}^W \end{pmatrix}$$

$$\delta F_{ij}^X = 1/2 \sum_{kl} a_{k,l}^X F_{i-k,j-l}$$

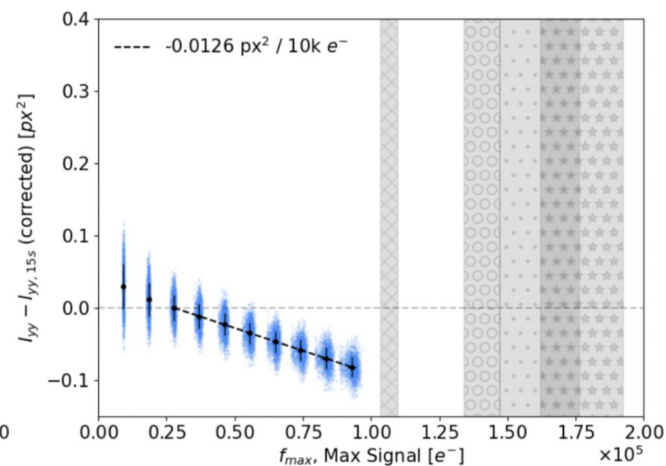
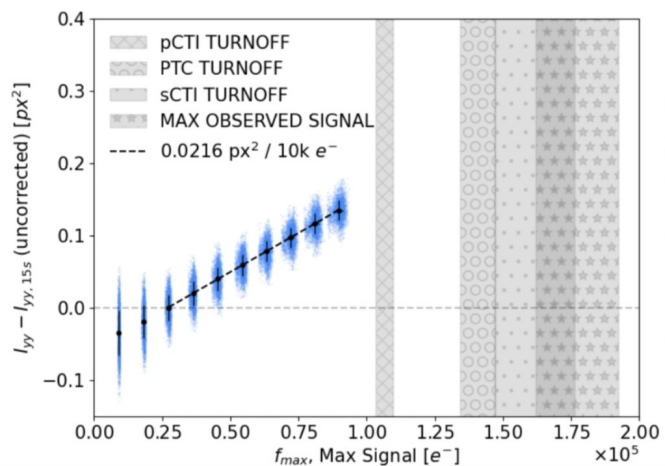


Corrects  
>95% of the  
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$I_{xx}$

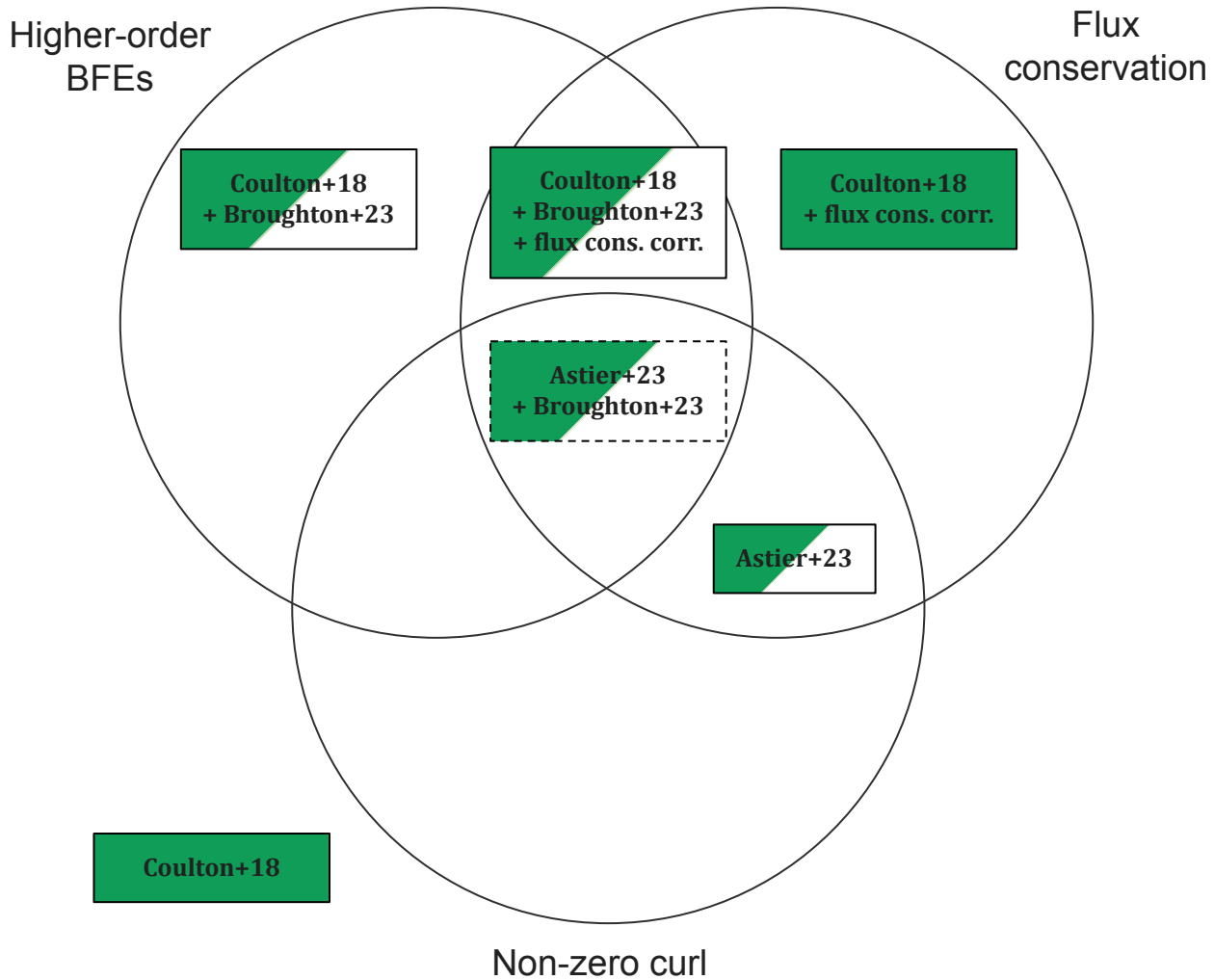


$I_{yy}$



# ISR Roadmap for BFE

 = implemented/  
not implemented



# ISR Roadmap for BFE

- Add flux sampling method from Broughton+23
- Add `elec_fit` from Astier+23 (found here: <https://gitlab.in2p3.fr/astier/bfptc>)
- Add parallel CTI turnoff calculation and store as curated dictionary
- Add option to `PhotonTransferCurveSolveTask` to set `maxSignalAdu` to:
  - Parallel CTI turnoff? Or  $x.x\%$  of PTC turnoff?
- Add optional higher-order shape statistics on sources (up to 4th order?)